GROUND-BASED GRAVITATIONAL-WAVE OBSERVATORIES



Gran Sasso Science Institute INFN/LNGS and INAF









A new window into the Universe









Strain sensitivities as a function of frequency



Abbott et al. 2020, LRR

Observing run timeline and BNS sensitivity evolution





O5 volume = 15*O3 volume

Abbott et al. 2020, LRR

Radioactively powered transients



What is the future of the GW astrophysics?

The European 3G concept



Europe we developed the idea of a 3G GW observatory

- Factor 10 better (x1000 Volume) than 2G detectors
- Wide frequency, with special attention to low frequency (few HZ)
- Capable to work alone (but aiming to be in a 3G network)

. . .

• 50-years lifetime of the infrastructure



Triangular shape Arms: 3 → 10 km Underground Cryogenic increase laser power Xylophone

ESFRI proposal submitted in September

3G effort worldwide



NSF funded in 2018 the Conceptual Design Study of a 3G facility: Cosmic Explorer: 40km – L shaped detector

EXPECTED SENSITIVITY



The ET exquisite sensitivity and wide frequency band will make it possible:

• Large distances back to the early Universe



Detection horizon for black-hole binaries



The ET exquisite sensitivity and wide frequency band will make it possible:

- Large distances back to the early Universe
- access unexplored mass up to 10³ Mo



courtesy Colpi and Mangiagli

The ET exquisite sensitivity and wide frequency band will make it possible:

- Large distances back to the early Universe
- access unexplored mass up to 10³ Mo
- benefit of exceptional parameter estimation accuracy for very high SNR events



3G Science case WP

ET will provide a wealth of data that have the potential of triggering revolutions in astrophysics, cosmology and fundamental physics

A summary of the Science of ET

(see Maggiore et al. 2020)

Astrophysics

- Black hole properties
- origin (stellar vs. primordial)
- evolution, demography
- Neutron star properties
- interior structure (QCD at ultra-high densities, exotic states of matter)
- demography
- Multi-band and -messenger astronomy
- joint GW/EM observations (GRB, kilonova,...)
- multiband GW detection (LISA)
- neutrinos
- Detection of new astrophysical sources
- core collapse supernovae
- isolated neutron stars
- stochastic background of astrophysical origin



Fundamental physics and cosmology

- The nature of compact objects
- near-horizon physics
- tests of no-hair theorem
- exotic compact objects
- Tests of General Relativity
- post-Newtonian expansion
- strong field regime
- Dark matter
- primordial BHs
- axion clouds, dark matter accreting on compact objects
- Dark energy and modifications of gravity on cosmological scales
- DE equation of state
- modified GW propagation
- Stochastic backgrounds of cosmological origin



THE UNEXPECTED...ET will be a "discovery observatory"!

Multi-messenger in the ET era

Binary systems of Compact Objects



Large increase of detection rate

Better parameter estimation

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ET capabilities

Astrophysical simulations for BNS from population synthesys code



MERGER RATE EVOLUTION impact of SFR and metallicity uncertainty

Santoliquido et al. 2020,



ET DETECTION EFFICIENCY



ALL ORIENTATION



EINSTEIN TELESCOPE DETECTION/SKY LOCALIZATION up to z=0.26

ΕT



ET+LIGO/Virgo/KAGRA/LIGOindia



ET SNR>12 and LKVI included when SNR > 4

1 year of observations

Up to z=0.26



- Among ~4000 mergers per year detected 3000 per year
- For ET 100 per year have sky loc < 10 sq. degrees
 - For ET+LVKI 1000 per year have sky loc < 10 sq. degrees
- For ET+LVKI 100 per year have sky loc < 1 sq. degrees

EINSTEIN TELESCOPEDETECTION/SKY LOCALIZATION up to z=1.8 =T ET+CE



1 week of observations

For ET+ CE 100 per week have sky loc < 10 sq. degrees



At z larger than 0.2 sky-localization from GRBs!



Thermal and non thermal emission components associated with BNS and NSBH merger

Ascenzi et al. 2020 arXiv:2011.04001





Ghirlanda

THERMAL EMISSION - KILONOVAE

OPTICAL BAND



- Too faint counterpart
- Large sky-localization/many contaminants



Joint detections for ET limited by optical instruments capabilities!!

Kilonovae detectable by the Vera Rubin Observatory survey up to 1 Gpc

In this volume

- ET about 100 event per year have sky loc < 10 sq. degrees
- For ET+LVKI 10³ per year have sky loc < 10 sq. degrees

A few tens to a few hundreds joint detections!

Three epochs of VRO 300s ToO observations in two filters (as in Chen et al. 2020)





HIGH-ENERGY



- GRB detectable up to high z
- Small number of contaminants
- Promising wide FoV hard-soft Xray instruments
- Good sky localization to drive a prompt EM follow-up





See Giulia Stratta talk

ET, ET+CE, ET+CE+CE







THESEUS in Multi-Messeger context

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V Sky and



GW detectors	THESEUS+GW detectors plausible joint observation time	aligned short GRB+GW detections	aligned & misaligned short GRB+GW detections
2G network z<0.107 (500 Mpc)	3.45 yr	~0.04	1.8
ET	1 yr (3.45 yr)	5.6 (19.2)	13 (46)
ET+CE	1 yr (3.45 yr)	7.4 (25.7)	16 (55)
ET+2CE	1 yr (3.45 yr)	8.7 (30.1)	18 (61)

Credit THESEUS Yellow Book

OTHER PROMISING HIGH-ENERGY COUNTERPARTS for SXI (0.5- 5 KeV sky loc < 1-2'





Magnetar? or

High latitude emission from structured jet?





What happen off-axis?



Ascenzi et al. 2020 A&A

Promising X-ray couterparts!

THESEUS AND ET as multi-probes of the early Universe

Disentangle astrophysical PoPIII from primordial BHs



De Luca et al. 2102.03809

Any BBH merger at z>30 will be of primordial origin

- Difference between ET and CE due to the better ET sensitivity at low frequencies
- Note: accurate measurement of z is also needed !

Cosmology and dark energy with ET

Modified GW propagation

Coalescing binaries measure a ``GW luminosity distance" different from the standard (electromagnetic) luminosity distance !

PARAMETRIZATION OF MODIFIED GW PROPAGATION

$$\frac{d_L^{\,\rm gw}(z)}{d_L^{\,\rm em}(z)} = \Xi_0 + \frac{1 - \Xi_0}{(1 + z)^n}$$

This parametrization is very natural, and fits the result of (almost) all modified gravity models

Belgacem et al.

PRD 2018, 1712.08108 PRD 2018, 1805.08731 JCAP 2019, 1907.02047

• Standard sirens in Giulia Stratta talk

Ratio of the gravitational to electromagnetic luminosity distance in a modified gravity model



Modified gravity models: the counterpart could be at a different distance with respect to luminosity distance from the GWs!

GW/EM the same travel time \rightarrow temporal EM/GW coincidence

Belgacem et al. JCAP 2019, 1907.02047



- JOINT GW/GRB detections
- Small error in dL^{GW}
- High z





Belgacem et al. JCAP 2019, 1907.02047



→ Large effect could be detectable even with just a single standard siren at ET

Few hundreds of joint GW-GRB detection $\rightarrow \Delta \Xi_0 / \Xi_0 \approx 1\%$ or better





⇒ GWs could become the best experiments for studying dark energy!





Expected brilliant synergy!