

Mrk 110 multiwavelength variability and prospects for THESEUS

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(in collaboration with
I. Mc Hardy, E. Cackett, A. Baarth, K. Horne and many others)

AGN powered by SMBH at center of galaxies

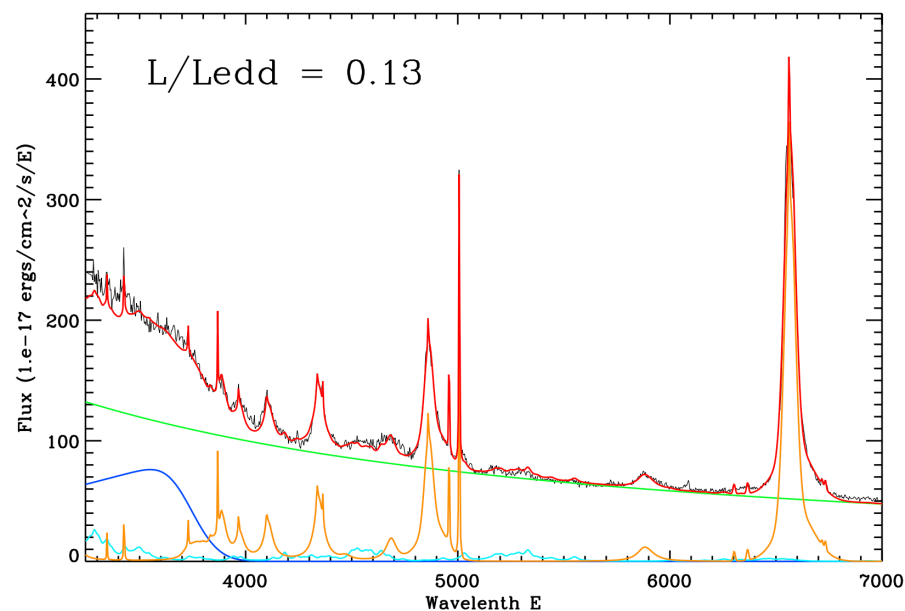
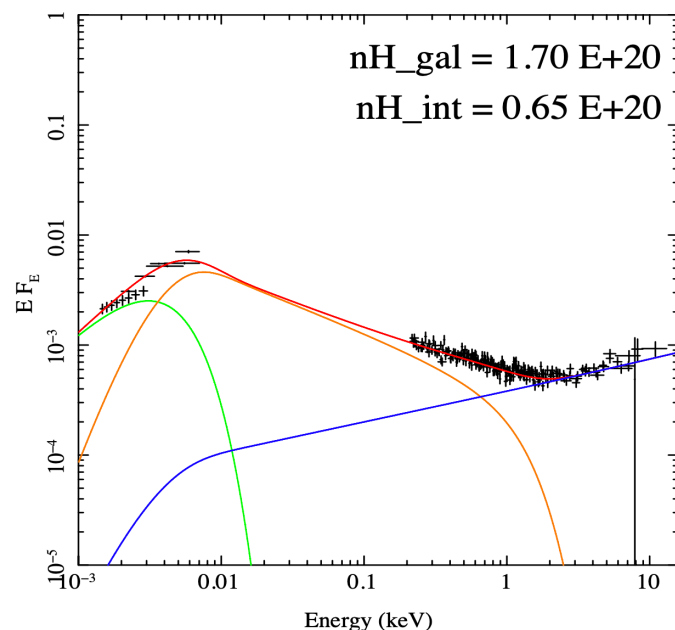
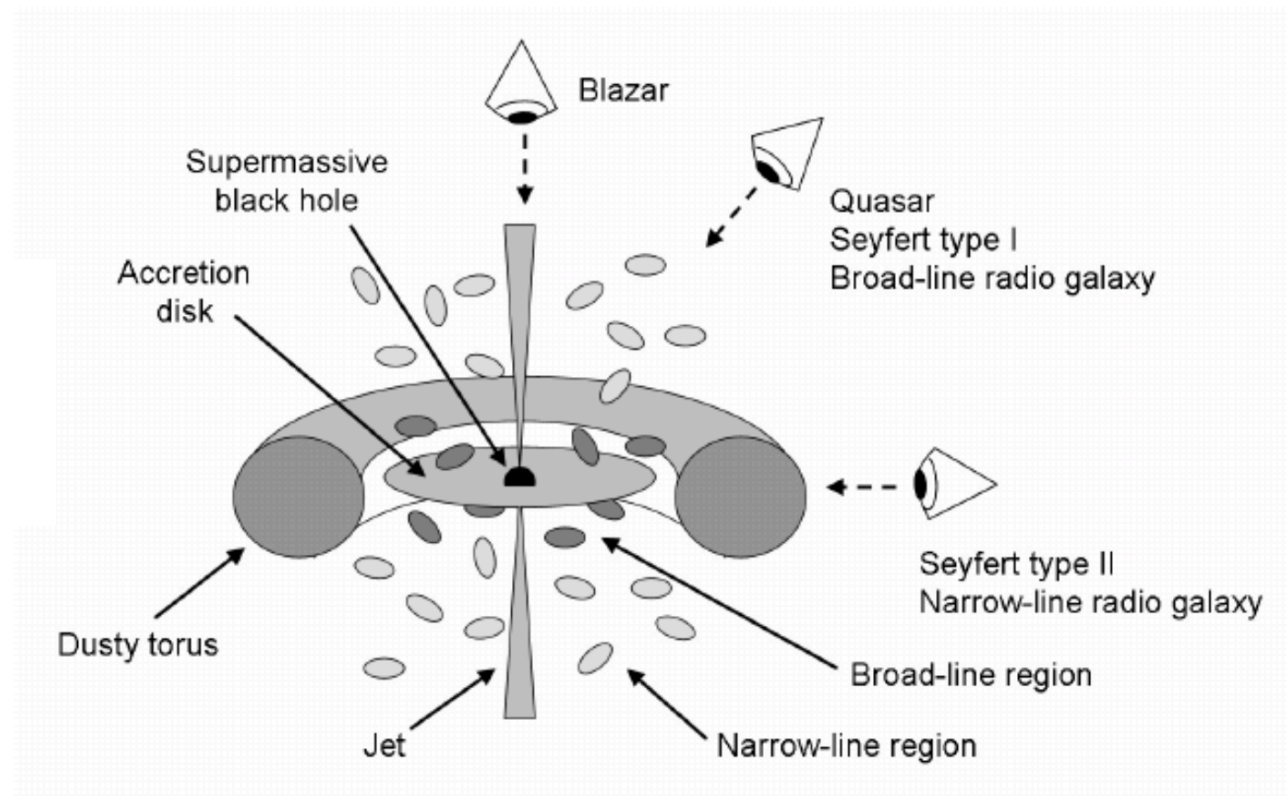
Complex accretion flow, with multiple components (dominating at different energies)

Corona: X-rays

Disk: optical UV

**Broad Line region: optical/UV lines
(possible contribution to continuum)**

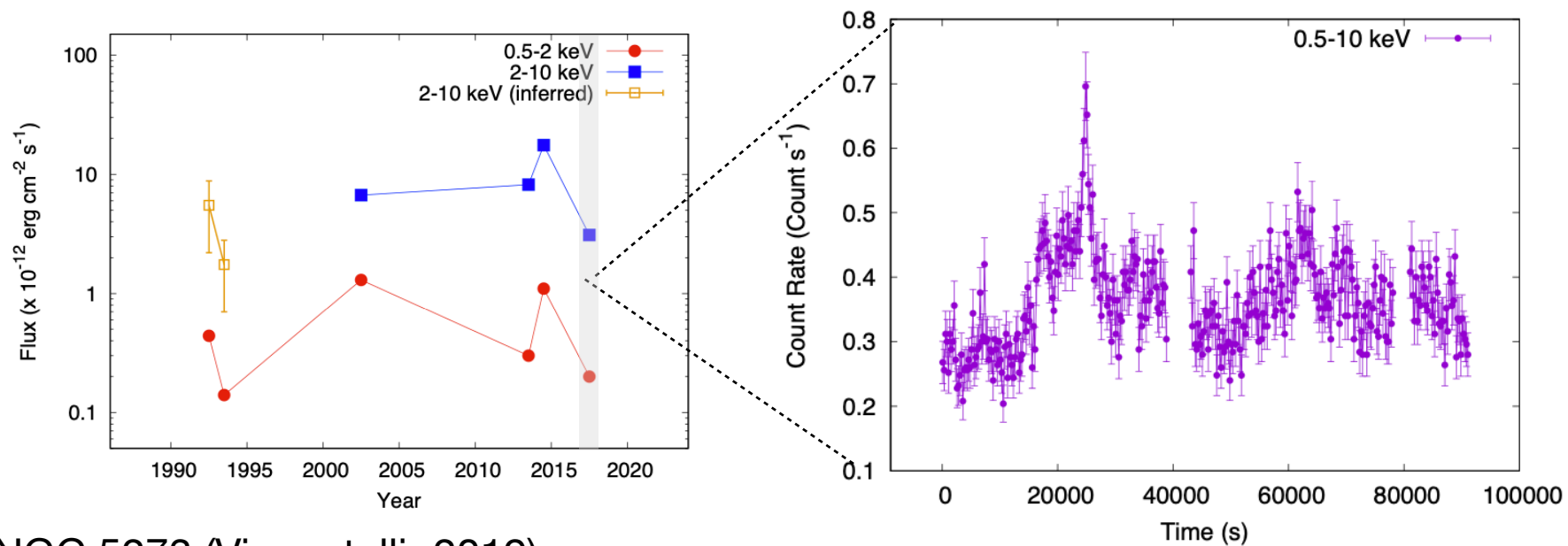
Torus: (IR)



Variability as a probe to constrain geometry

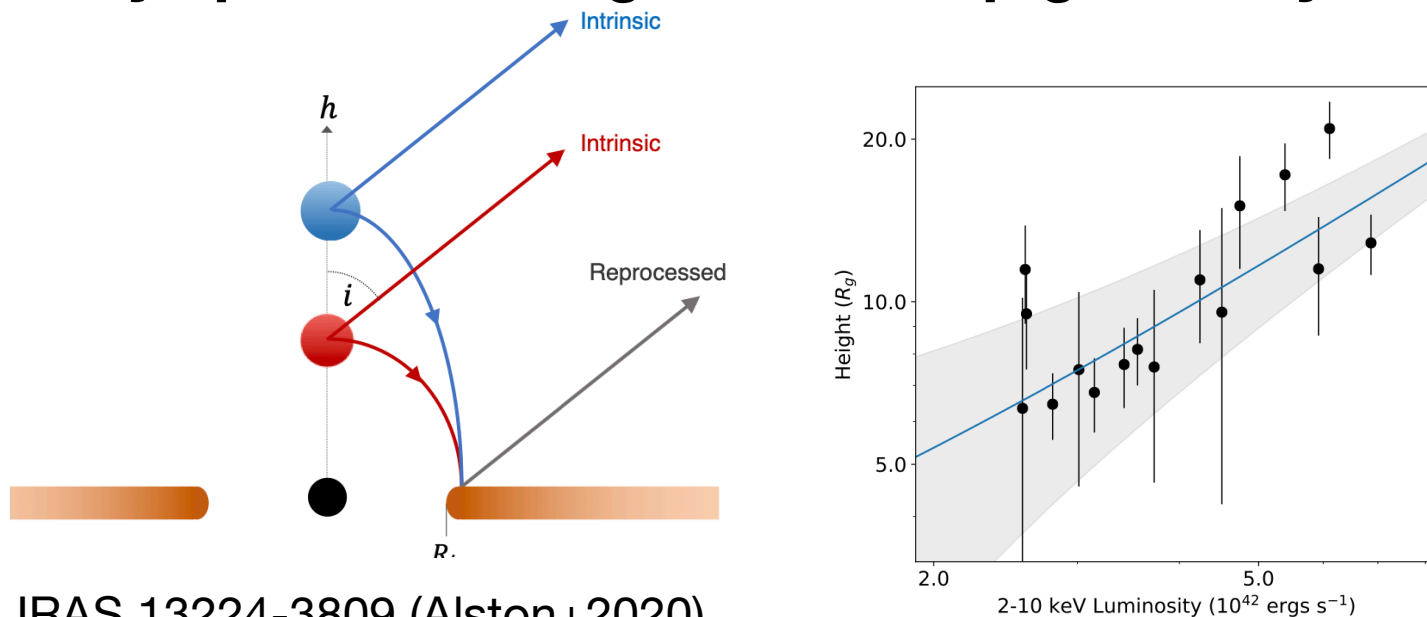
AGN are extremely variable in optical and X-rays

Timescales from ~ 100 s to years



NGC 5273 (Vincentelli+2019)

X-ray spectral-timing able to map geometry close to BH

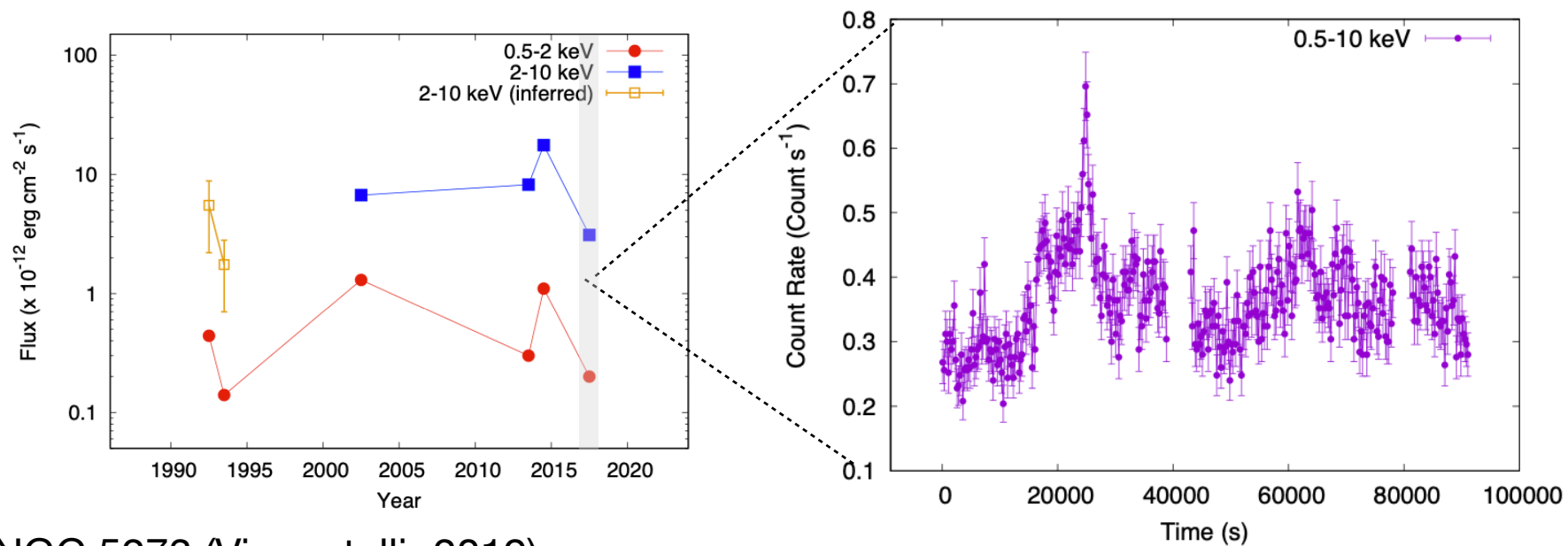


IRAS 13224-3809 (Alston+2020)

Variability as a probe to constrain geometry

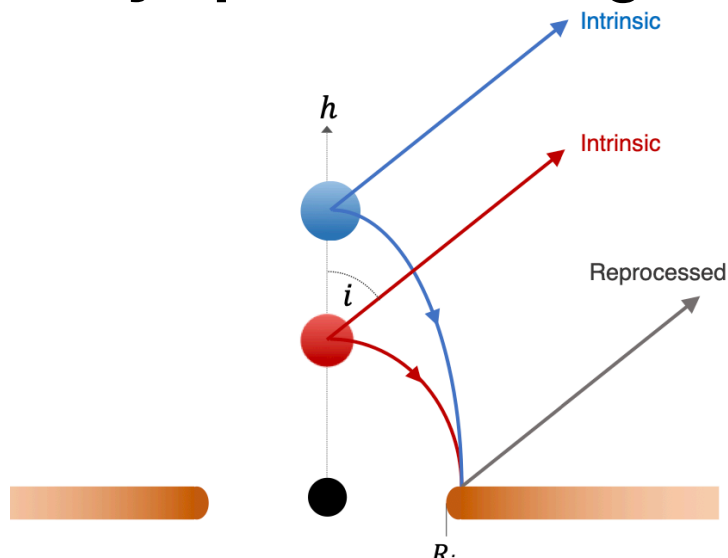
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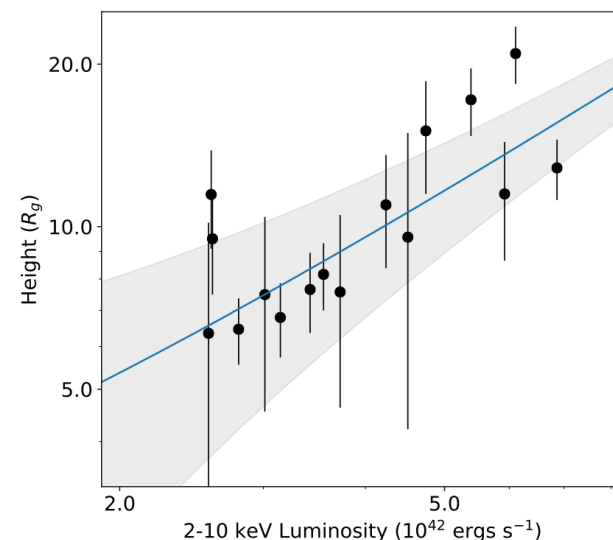


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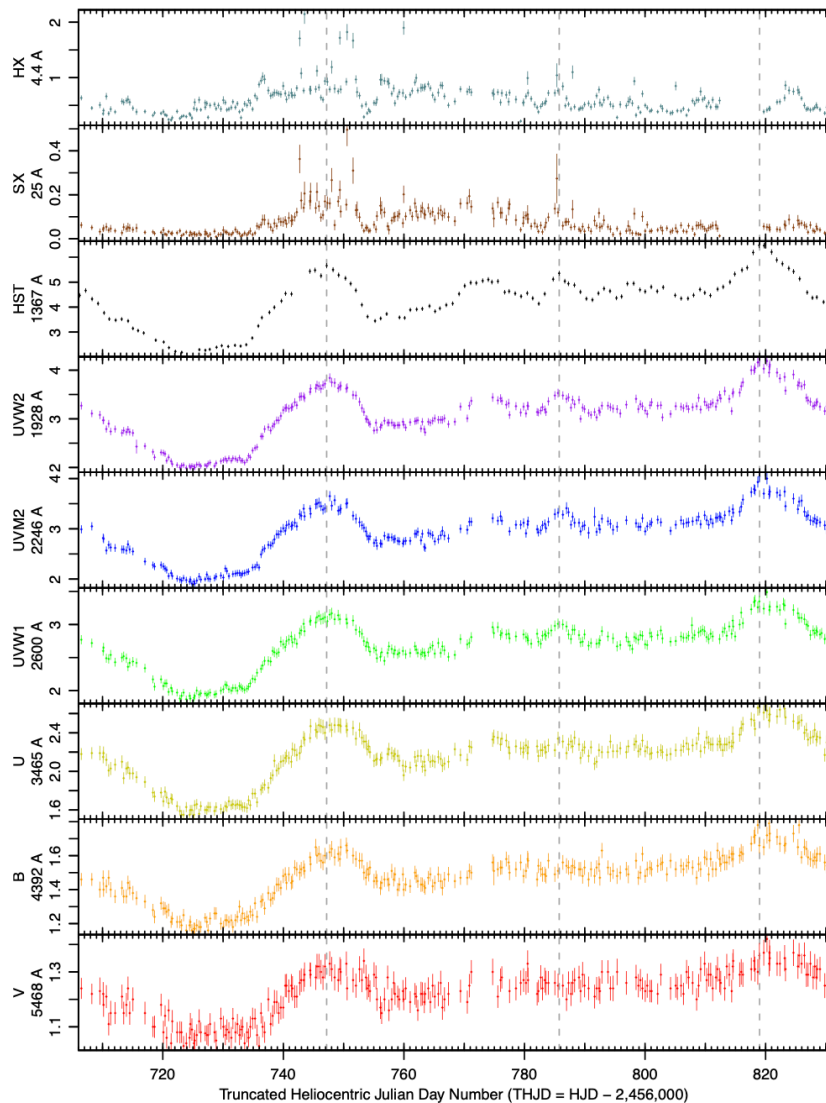
IRAS 13224-3809 (Alston+2020)



Same Experiment can be done with X-ray/Optical observations!! (outer disk)

Swift role for AGN variability

(NGC 5548, Edelson+2017)



Thanks to Swift schedule flexibility

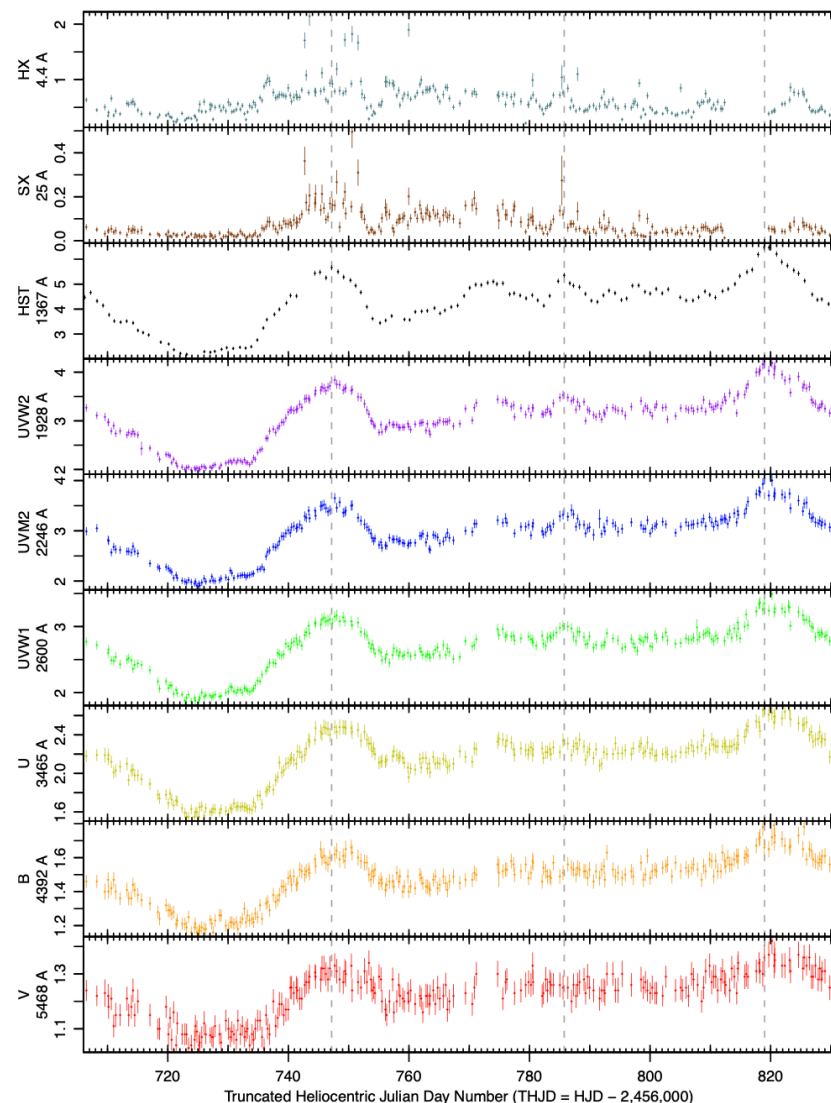
Excellent quality Optical-UV /X-ray monitoring data

~few observations per day for hundreds of days.

Very difficult to obtain from the ground

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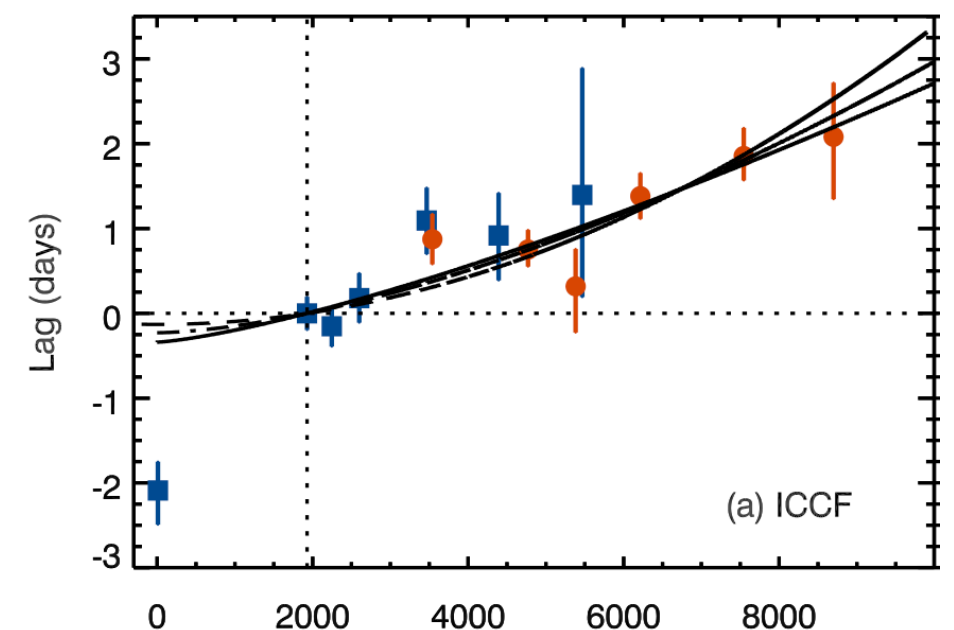
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(Mrk 142, Cackett+2018)



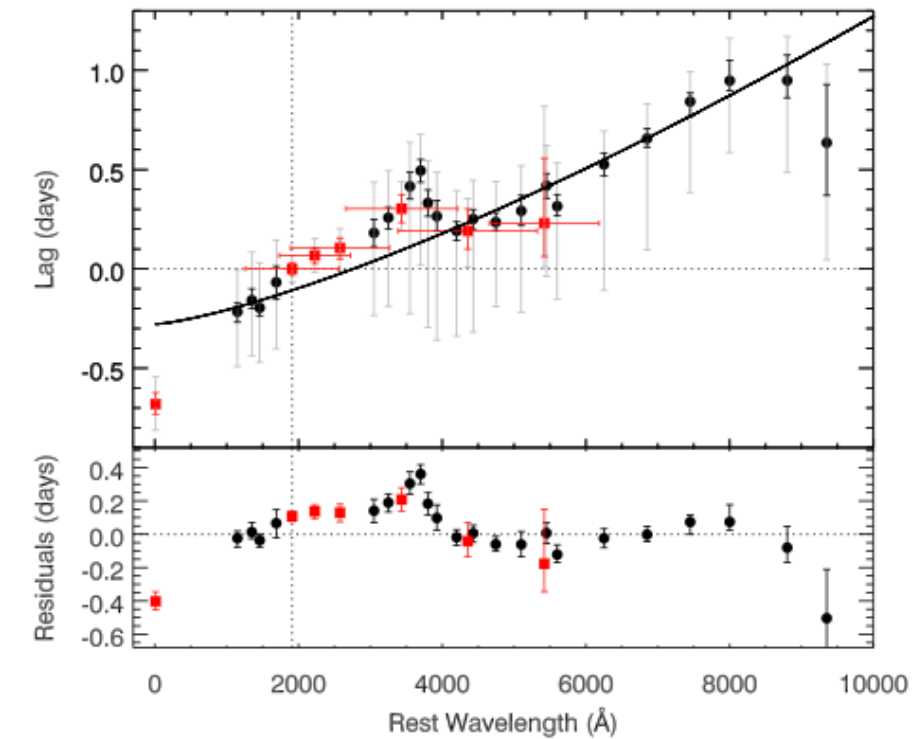
Study of LAGS increased our understanding of disk:

- not consistent with simple lamp-post model**
- uv/optical follow expectations, but x-ray do not**

Evidence for BLR also in the continuum

Spectroscopically resolved observations show extra component

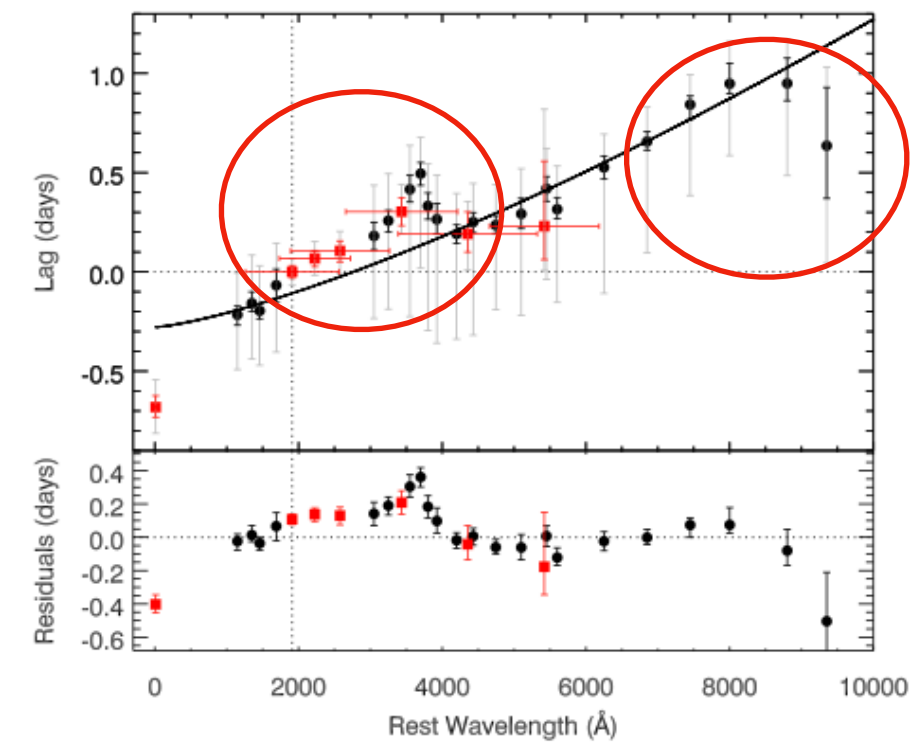
(NGC 4593, Cackett+2018)



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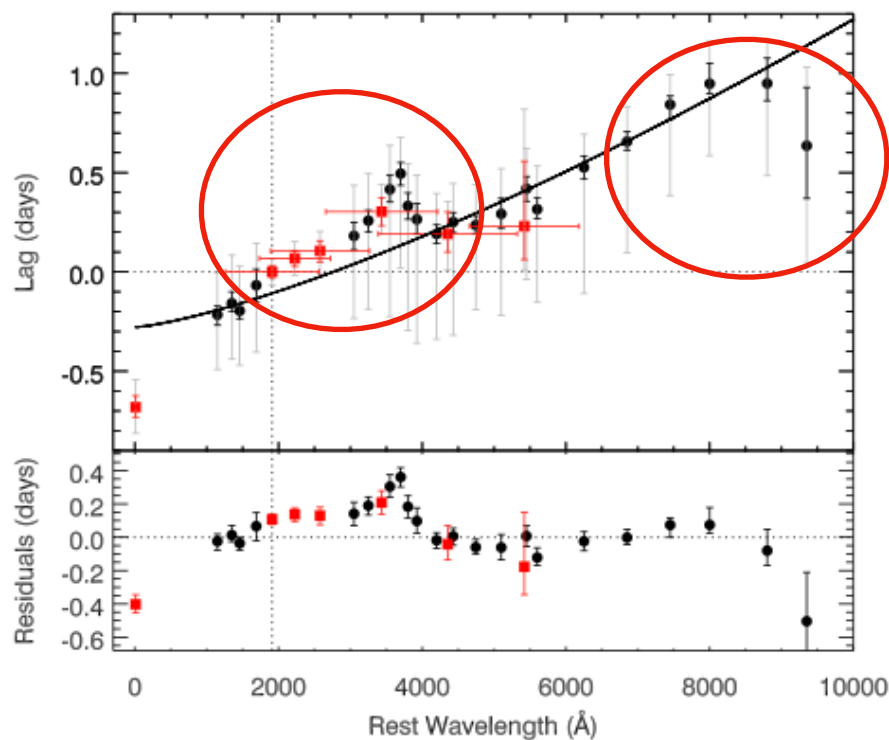
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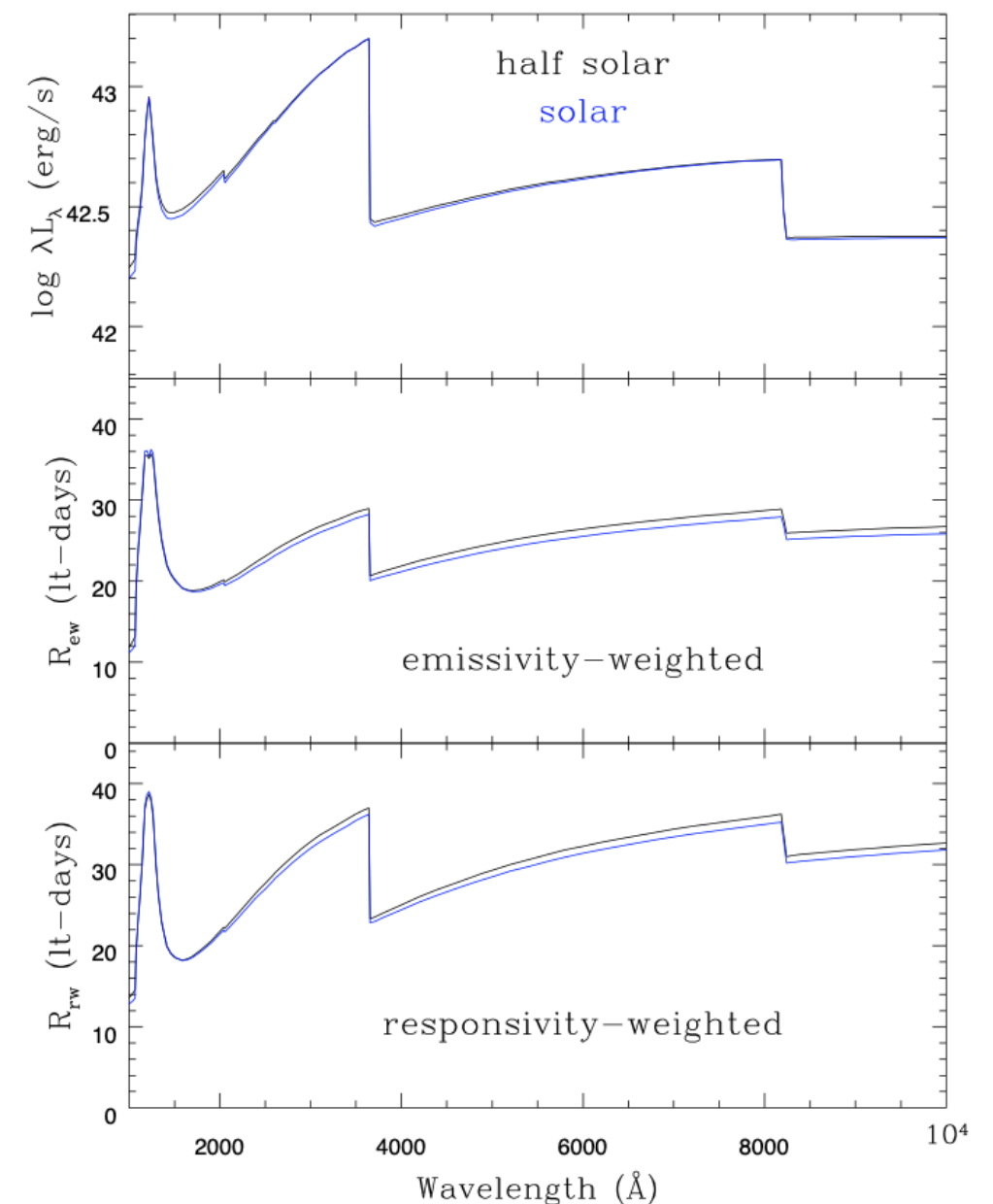
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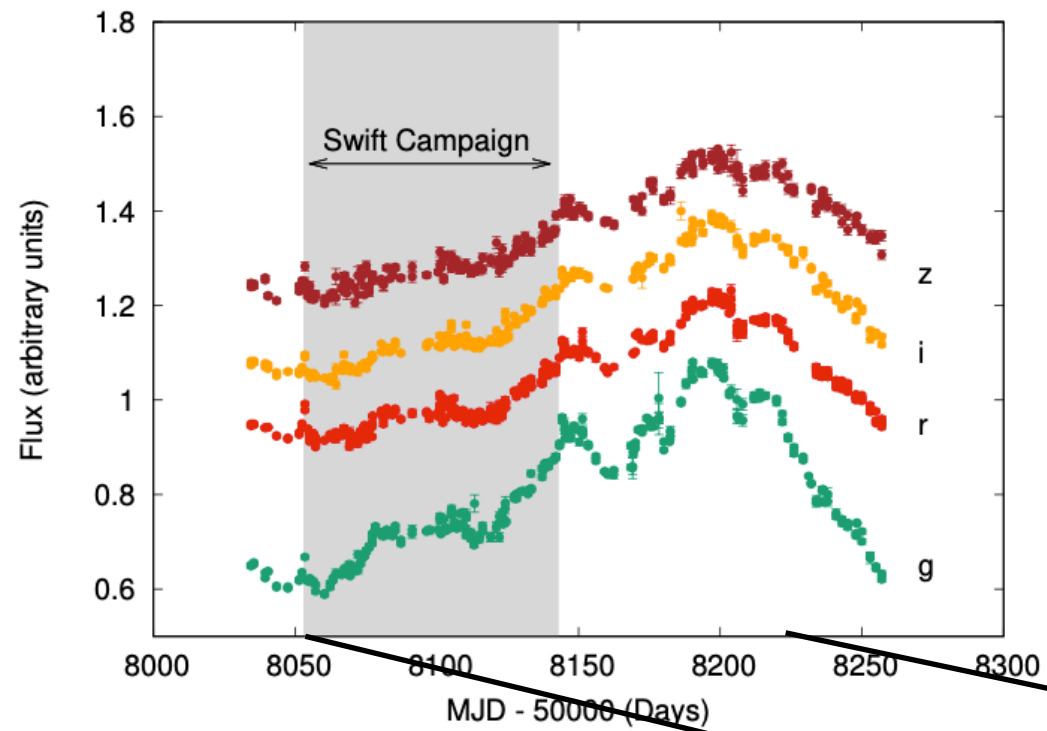
This can be explained in terms of BLR diffuse continuum!

Not only disk, but also BLR!

Simulations from Korista & Goad 2019

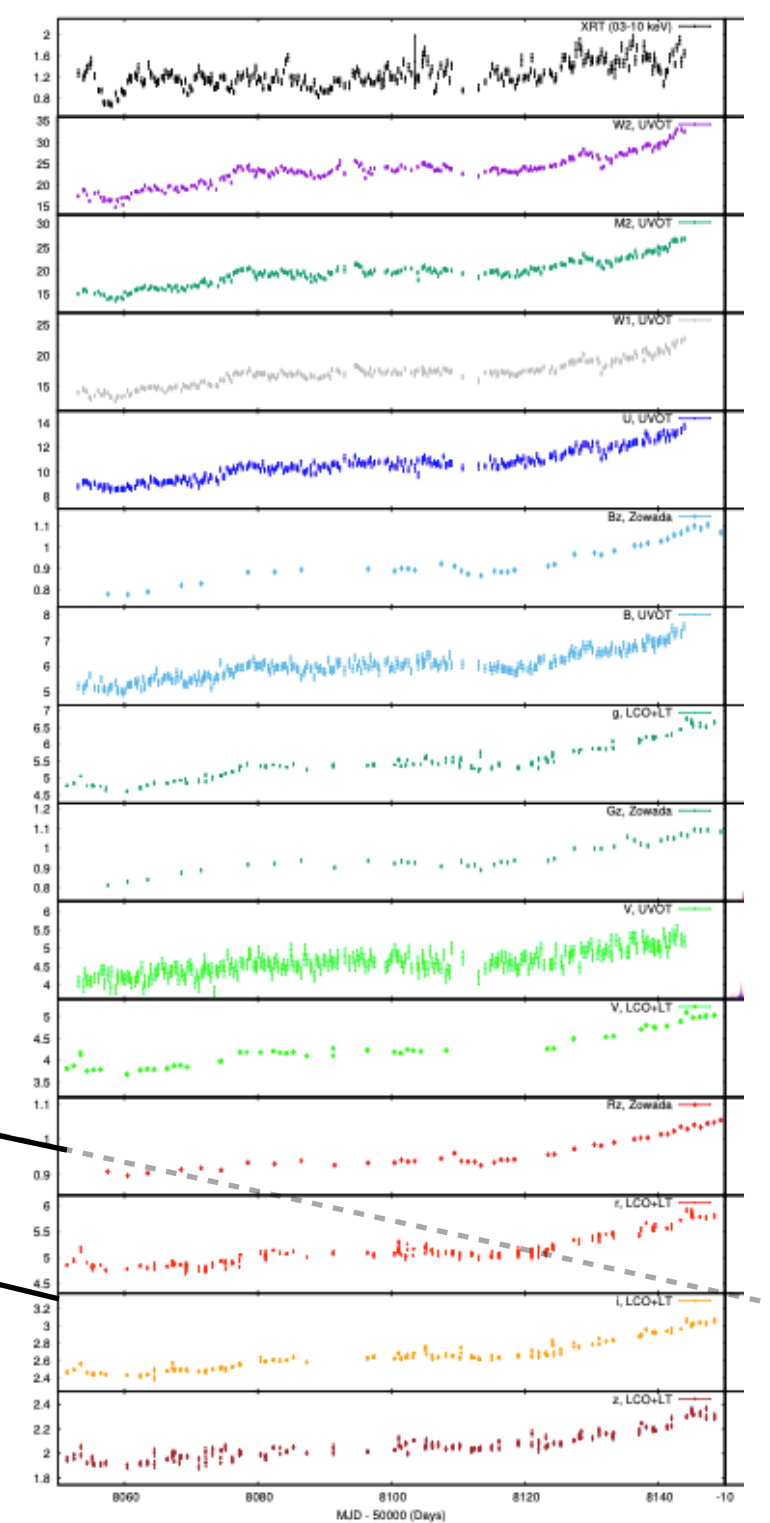


Mrk 110 campaign



>250 days campaign
including:

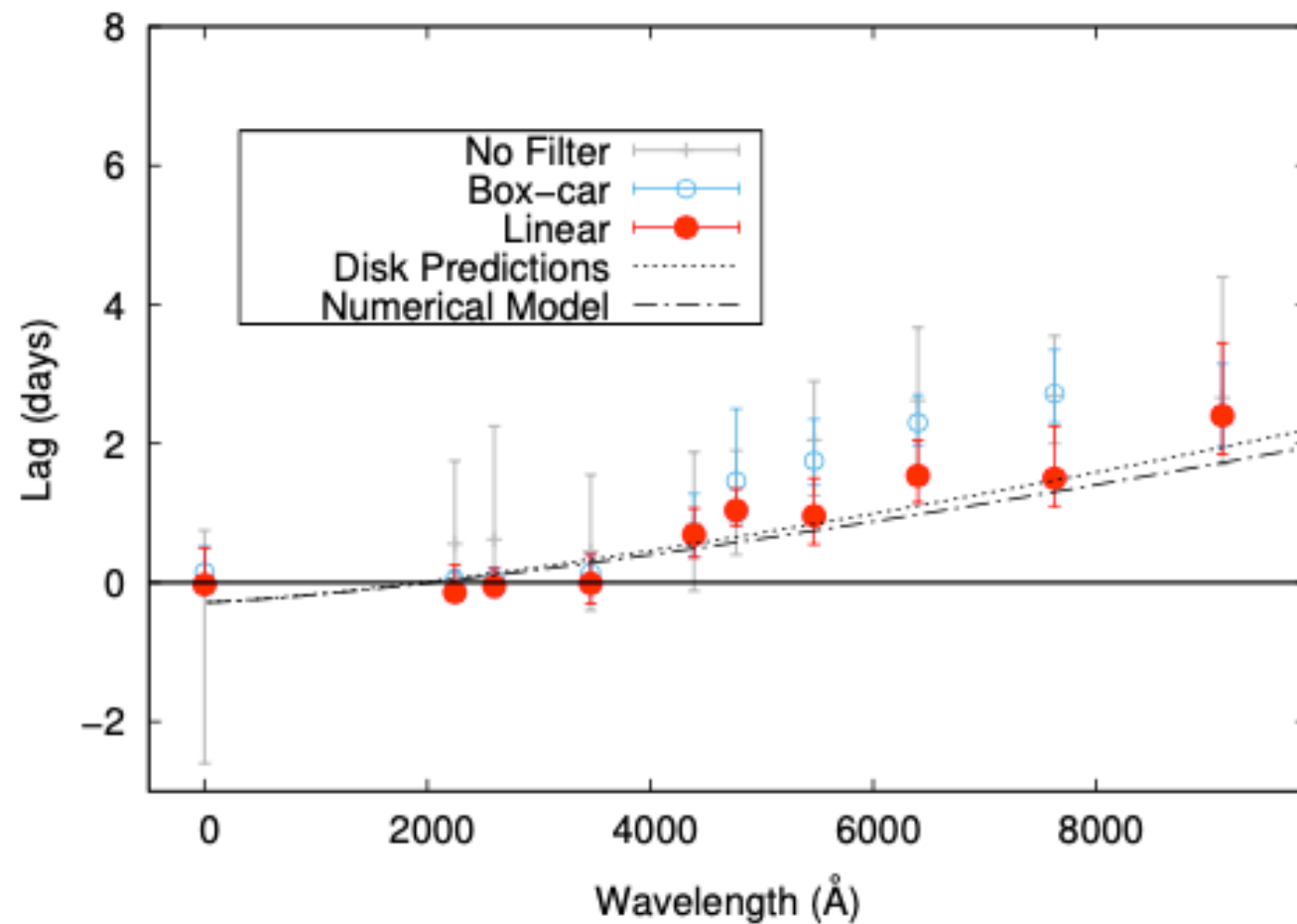
XRT
UVOT
LCO
LT



2 Clear Behaviours:

-Short term variation
-Long term trend

Short term lags



Lags vs UVW2 band (1928 ang.)

Long term trend present in data

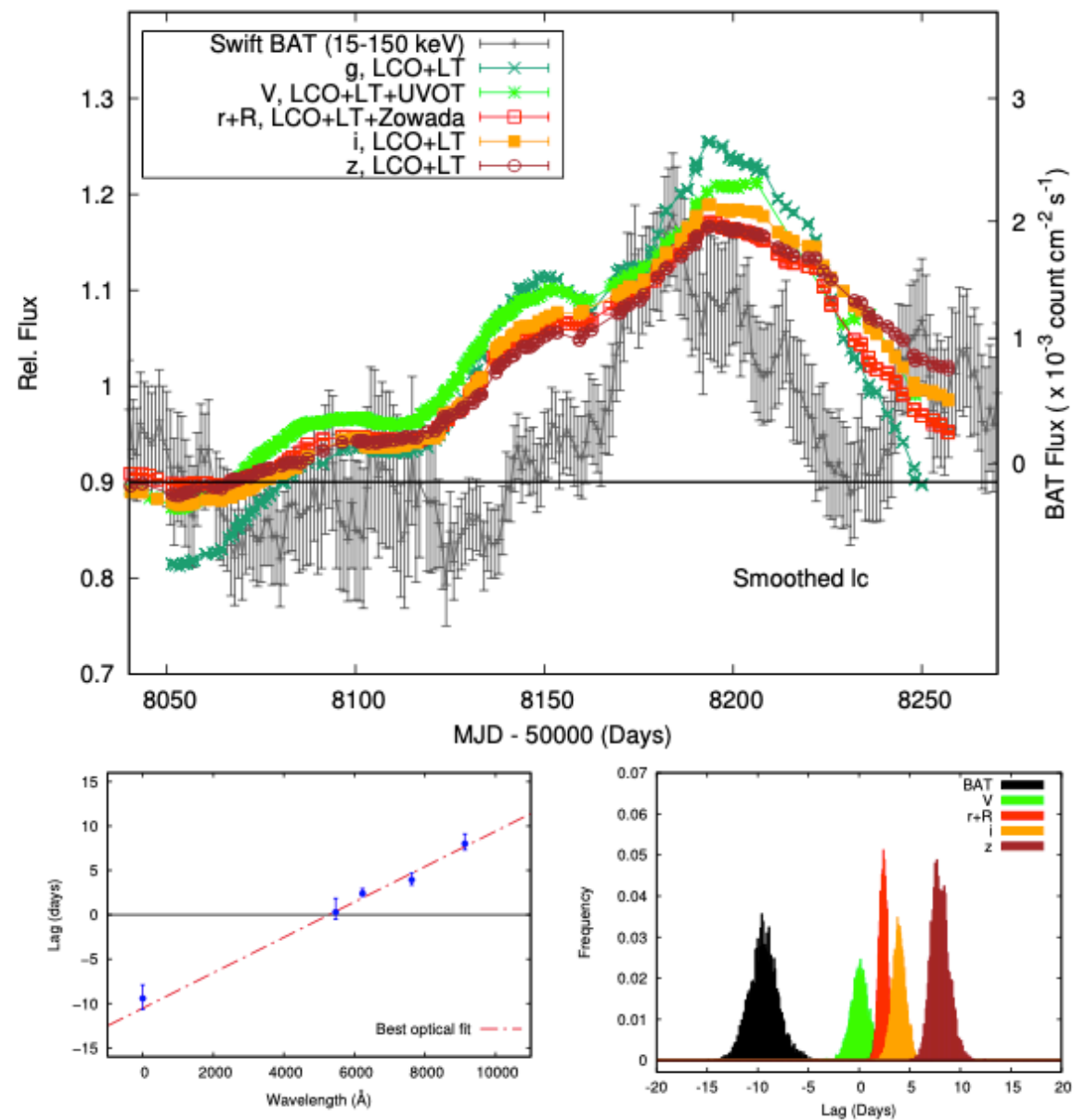
Different methods used to filter it

Lags consistent with disk predictions

ONLY WHEN FILTERING OUT LONG TIMESCALES

X-ray/UV consistent with 0, no UV excess

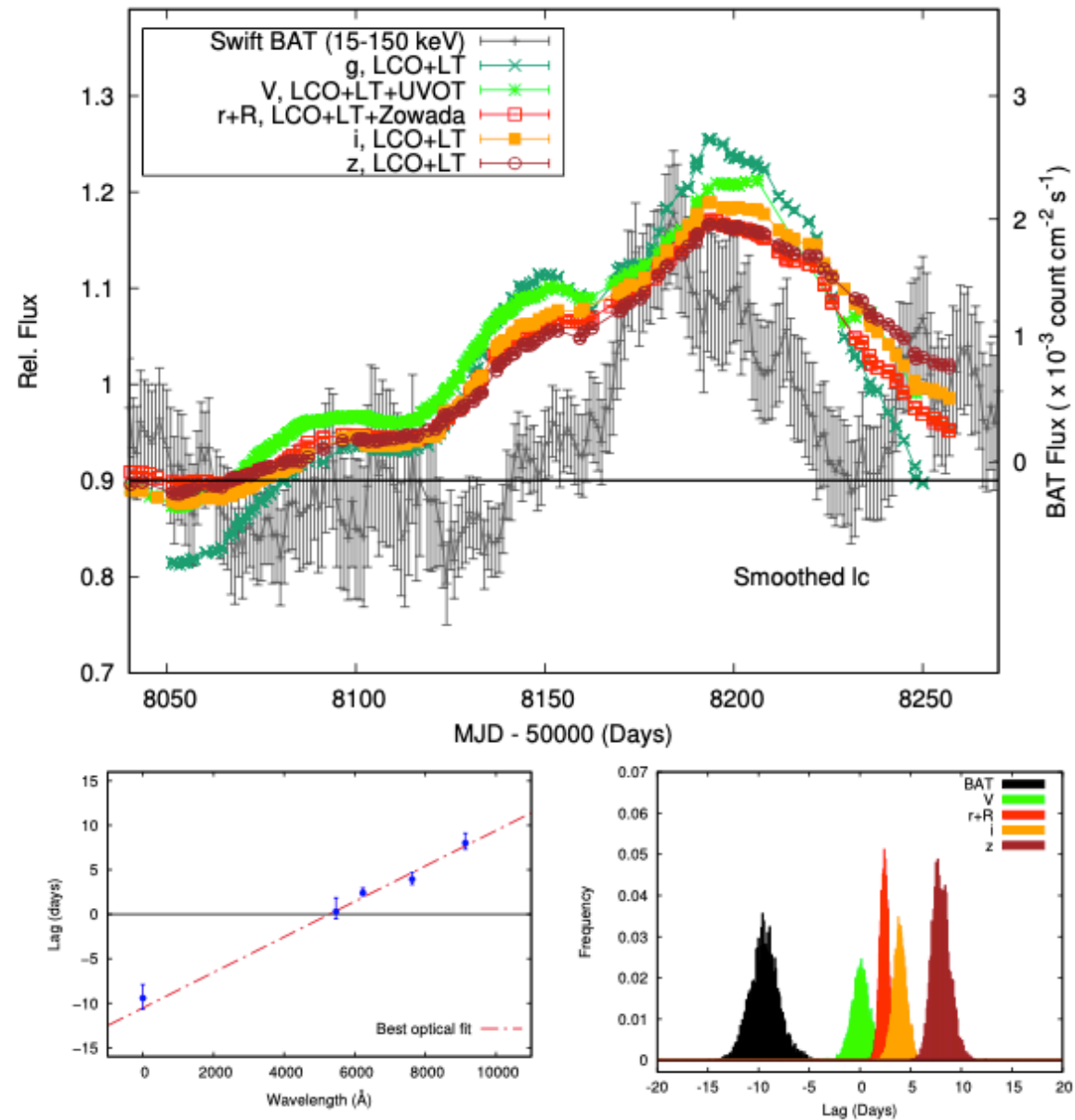
Long term trend



Longer variations show a much longer lag

Not consistent with disk predictions.

Long term trend



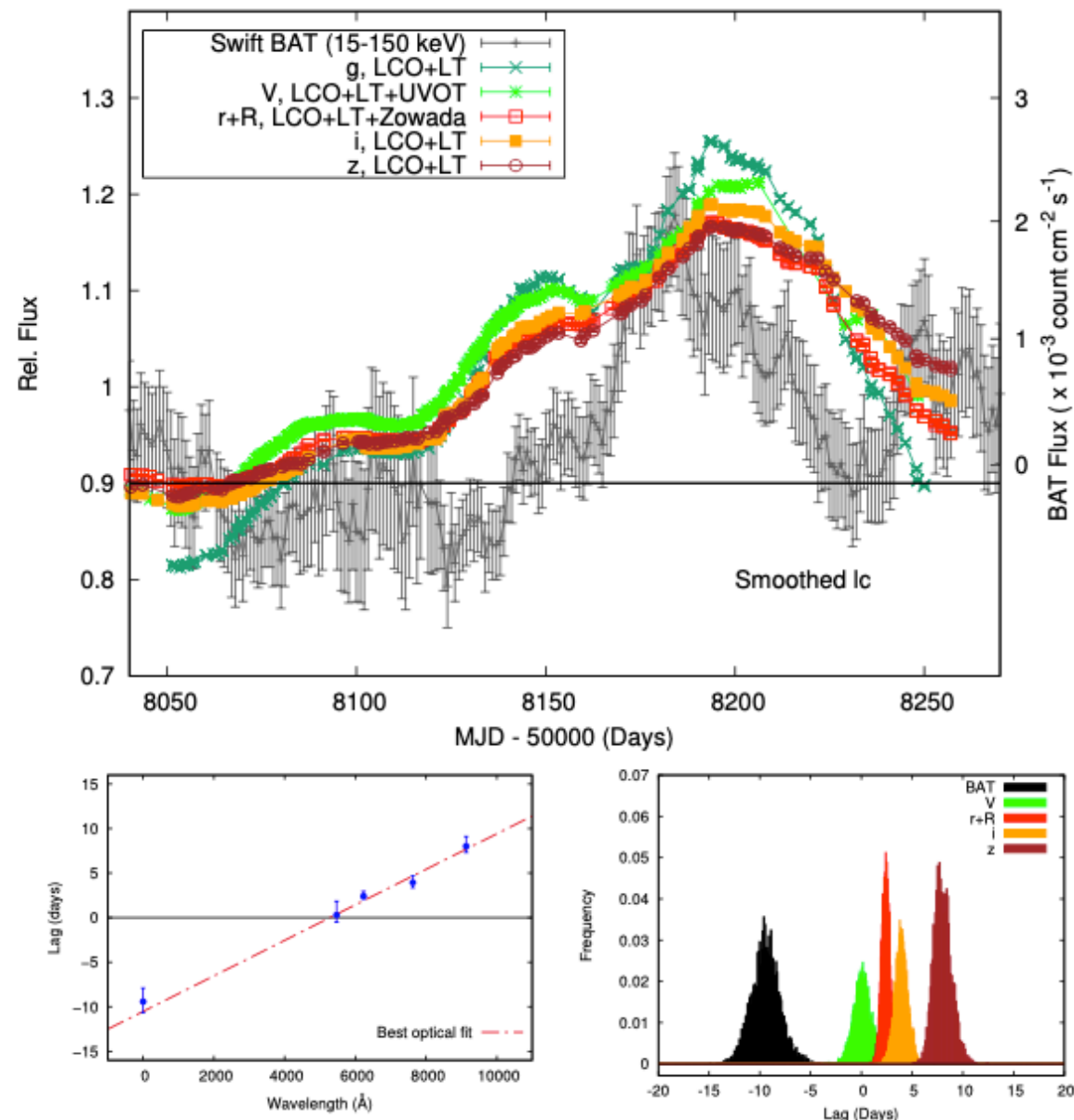
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SED evolves time, flare observed in BAT

This suggests stronger contribution of BLR

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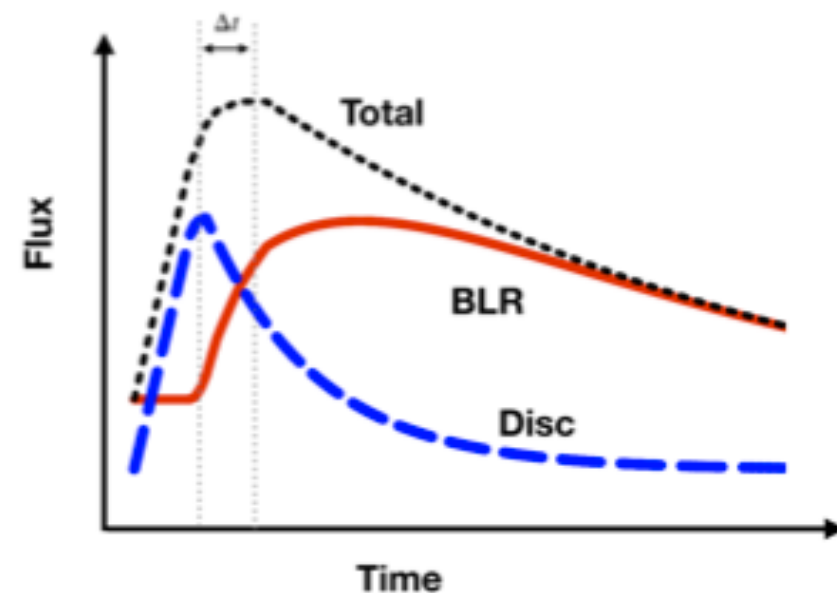
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Possible change in inner geometry
could trigger it

but no pointed observation
to test hypothesis
(perfect case for Theseus)



Why THESEUS can be crucial?

THESEUS is the evolution of Swift

Not only flexibility and sensitivity...but also extension to IR!

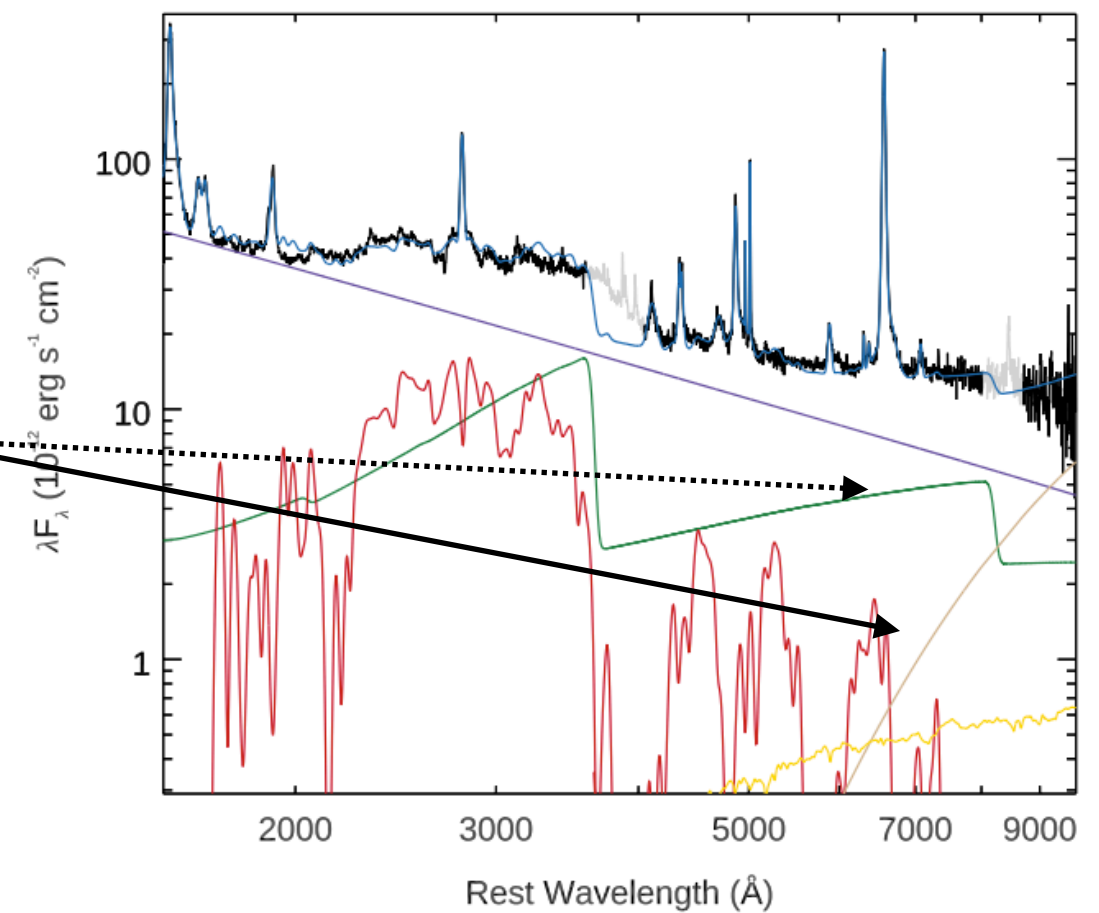
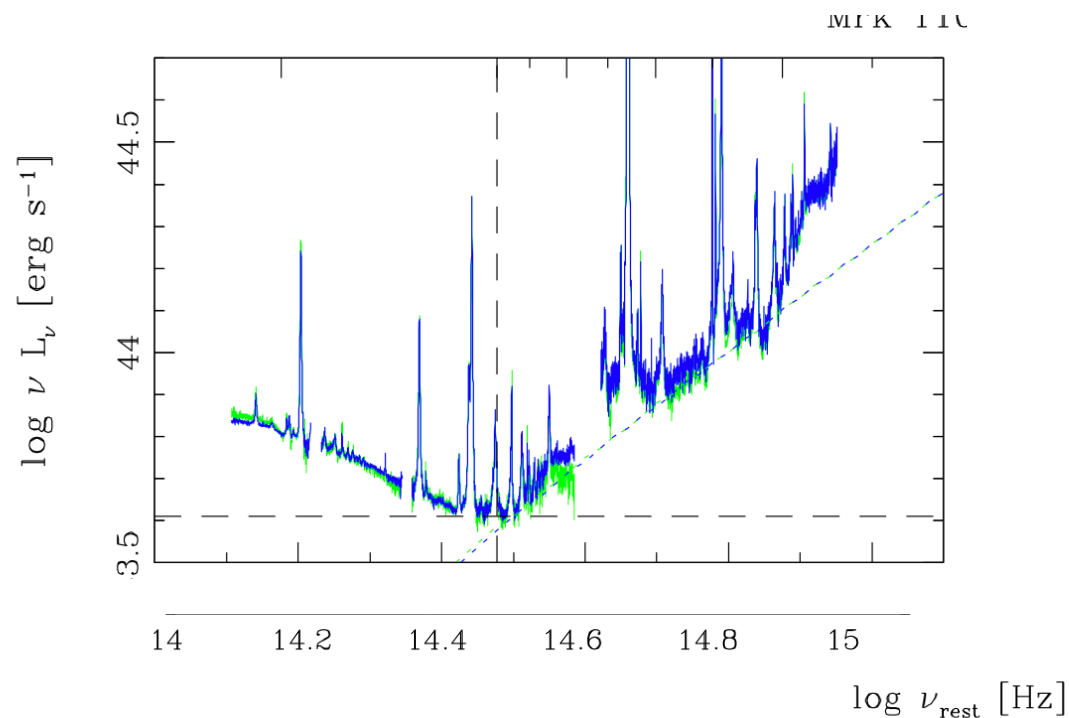
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Vincentelli+subm.

**Covering >800 nm can probe
BLR Diffuse continuum and Torus variability**



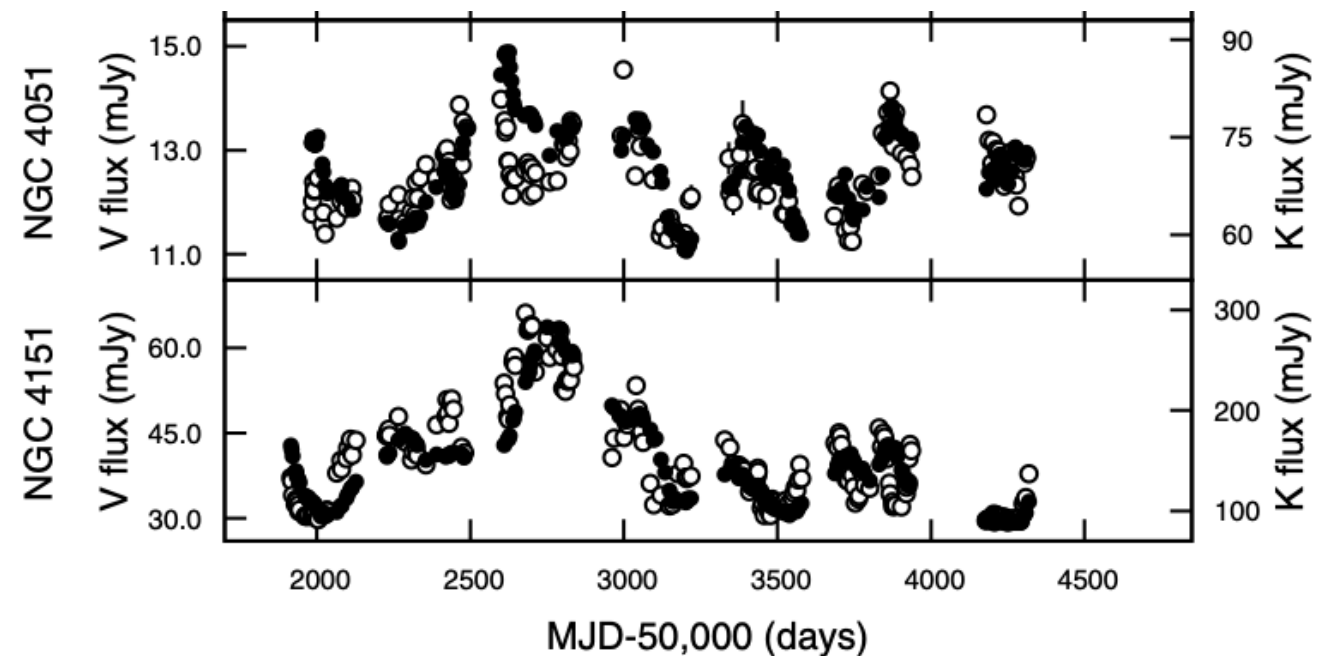
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**Longer wavelengths lightcurve data
are mainly from ground
very sparse sampling**



THESEUS (as Swift) is able to perform ~few obs. per day

probe with much higher detail the X-ray/IR (depending on the observing window)

CONCLUSIONS

Swift intensive monitoring of AGN is changing our view of AGN accretion flow

**A ground base+ Swift campaign of Mrk 110
revealed the presence of multiple components at different timescales**

**THESEUS thanks to its flexibility and multiwavelength coverage
can play a crucial role, especially for the BLR**

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THANK YOU